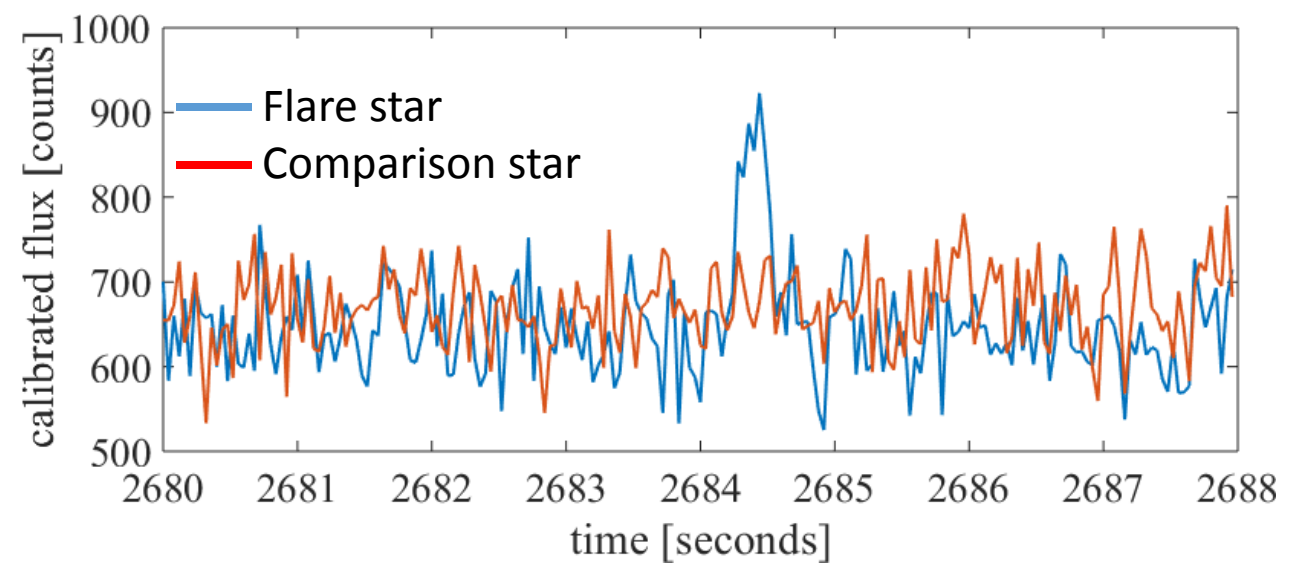


# A High Foreground of Sub-second Glints

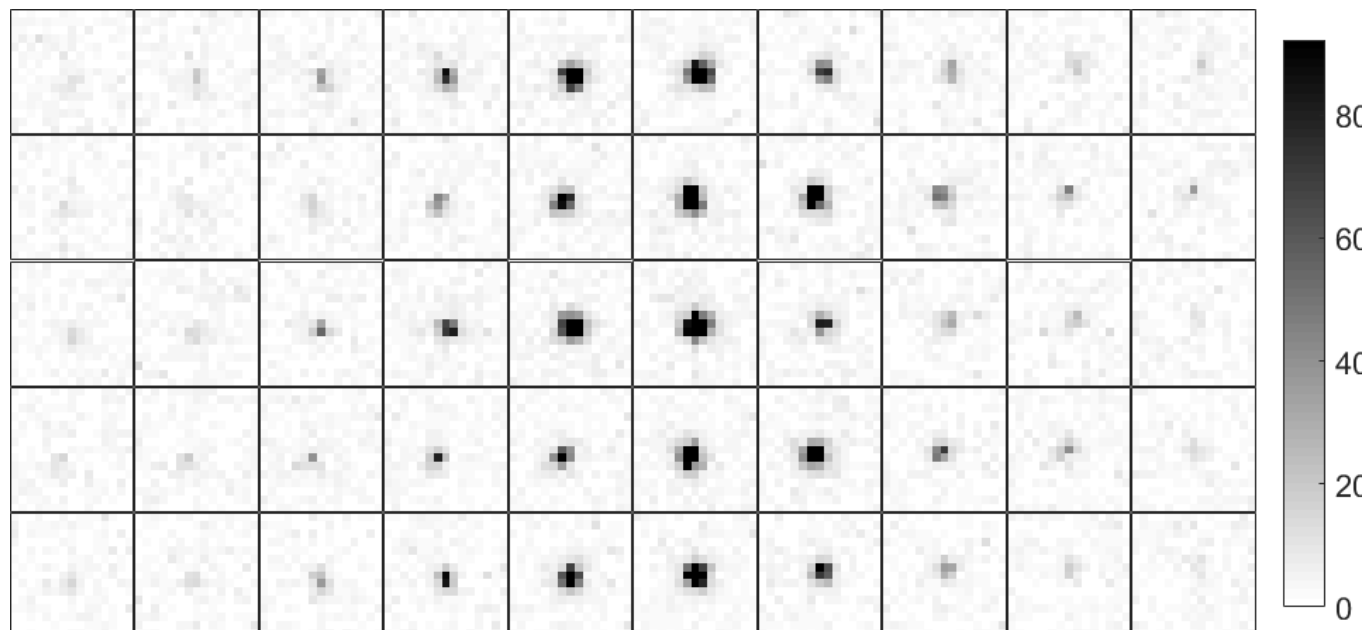
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<https://arxiv.org/abs/2011.03497>

**The sample.** We discovered point-like, sub-second duration flares of *unknown origin*. Some of the flares coincided with bright stars, some did not have apparent hosts. The magnitude of the flares was 9-11 (GAIA  $B_p$  equivalent). The duration of the flares was 0.05-0.3s.



**Fig 1.** An example lightcurve for two stars in the same field, with the blue curve showing a typical flare. The star brightens by  $\sim 50\%$  in less than a second.

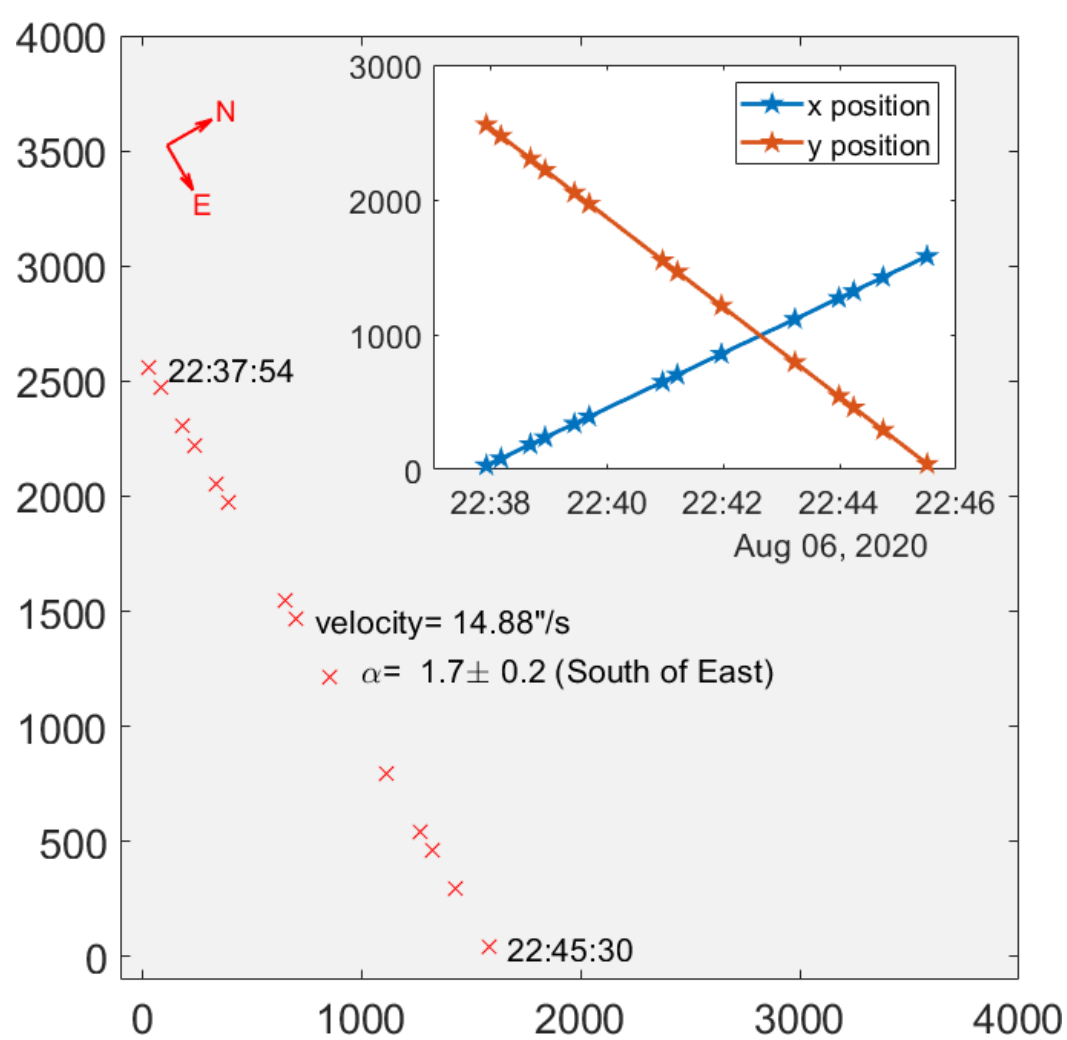


**Fig 2.** Each row is a single flare event, each square an individual image taken 40ms apart. Flares were observed a few tens of seconds apart in different parts of the field.

**Possible sources** of sub-second optical transients:

FRB or GRB counterparts, dwarf-star flares, etc.

**However:** stars do not brighten by 50% in 0.2s, FRB counterparts are not so bright, GRB flashes are not so common.



**Fig 3.** The trajectory of a repeating flare across the field of view. The object, moving East at  $15''/s$ , is clearly a geosynchronous satellite.

**Nature of the flares.** After finding repeated flares moving in one image we can conclude these flares originate from geosynchronous/high-orbit satellites or space debris (old satellites, spent rockets, etc.).

**Physical properties.** Objects are tumbling with  $\sim$ minute periods. The reflector size is 1-10cm diameter.

**The rate** of flares close to the equatorial plane ( $\pm 20^\circ$ ) is 25-45 flares per day per  $\text{deg}^2$ .

**Conclusion.** Such glints can appear point-like. They are a common foreground for transient searches looking for short events ( $< 1$  exp. time), e.g., FRB/GRB counterparts.

**The observatory.** The *Weizmann Fast Astronomical Survey Telescope* (W-FAST) is a 55cm, f/2 telescope equipped with a fast sCMOS camera capable of taking continuous imaging of  $7 \text{ deg}^2$  at a frame rate of 25 Hz.

Current science done at W-FAST: a search for stellar occultations by Solar System objects (Kuiper Belt or Oort Cloud objects) and studying stellar variability at short time-scales.

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**Fig 4.** The W-FAST observatory in Mitzpe Ramon in the Israeli Negev.



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